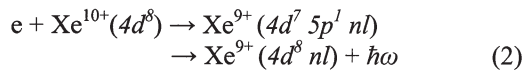


### §13. Satellite line of Xe<sup>9+</sup> Ions

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Recently EUV light sources from compact plasmas are now intensively studied for the next generation of lithography. The multicharged Xe ions emit EUV emission and are now investigated extensively. However we do not know the detailed atomic processes for there Xe ions. The satelliteline [1-2] transitions for parent emission of the types Xe<sup>10+</sup> (4d<sup>7</sup> 5p<sup>1</sup> → 4d<sup>8</sup>) ion are emitted by the radiative stabilization of doubly excited states in the Xe<sup>9+</sup> (4d<sup>7</sup> 5p<sup>1</sup> nl → 4d<sup>8</sup> nl). The doubly excited states are populated by the processes of dielectronic capture:



The dielectronic recombination processes is resonated by the incident electron energy, the required being equal to the energy of the doubly excited state relative to the ionization energy of the Xe<sup>9+</sup> ion. The autoionization takes place only when the energy of the doubly excited states is above the threshold energy.

The emissivity of a dielectronic satellite transition from the autoionizing state *i* to a final state *f* is proportional to a Qd (*i*<sub>0</sub>, *i*, *f*) factor which expresses the rate of dielectronic capture into the autoionizing state followed by radiative decay to the final state. Thus, the emissivity of a satellite lines produced by dielectronic recombination can be expressed as

$$I_s = n_e n_z \left( \frac{h^3}{2\pi m k_B T_e} \right)^{3/2} \frac{e^{-E_{i,i_0}/k_B T_e}}{2} \times \langle Qd(i_0) / g_0 \rangle \quad (3)$$

$$Qd(i_0, i, f) = \frac{g_i Aa(i, i_0) Ar(i \rightarrow f)}{\sum_{all i_0} Aa(i, i_0) + \sum_{all f'} Ar(i \rightarrow f')} \quad (4)$$

where *n<sub>e</sub>* is the electron density, *n<sub>z</sub>* is the ion density, *h* is the Plank constant, *k<sub>B</sub>* is the Boltzmann constant, *g<sub>i</sub>* and *g<sub>0</sub>* are the statistical weights of the autoionizing state (*i*) formed by dielectronic capture and the target state (*i*<sub>0</sub>) before dielectronic capture, *E<sub>i,i0</sub>* is the resonance energy, and <Qd(*i*<sub>0</sub>) / *g*<sub>0</sub>> represent the average value over all initial state the Xe<sup>10+</sup> (4d<sup>8</sup>).

To understand the relation of the satellite lines and the radiative decay lines of Xe<sup>10+</sup> ions, we need to calculation of the emissivity of a lines produced by collisional excitation rate coefficient from level *i*<sub>0</sub> to level *i'* using the Mewe empirical formula based on the Bethe approximation of parent ions can be expressed as

$$I_s = n_e n_z C_{i_0, i'}(T_e) \quad (5)$$

$$\begin{aligned} C_{i_0, i'}(T_e) = 3.69 \times 10^{-12} \frac{1}{\sqrt{T_e}} \frac{1}{E_{i_0, i'}} P(E_{i_0, i'} / T_e) e^{-E_{i_0, i'} / T_e} \\ \times \langle Ar_{i', i_0} g_{i'} / g_0 \rangle \end{aligned} \quad (6)$$

where *g<sub>i'</sub>* and *g<sub>0</sub>* are the statistical weights of the excited state *i'* and ground states *i*<sub>0</sub> of Xe<sup>10+</sup>, *E<sub>i',i0</sub>* is the energy difference between *i*<sub>0</sub> level to *i'* level and <Ar<sub>*i',i0*</sub> \* *g<sub>i'</sub>* / *g*<sub>0</sub>> represent the average value over all initial state the Xe<sup>10+</sup> (4d<sup>8</sup>). Figure 1 shows the sum of the intensity factor (<Qd(*i*<sub>0</sub>) / *g*<sub>0</sub>>) for the satellite transitions from states of the 4d<sup>7</sup> 5p<sup>1</sup> 5l to 4d<sup>8</sup> 5l of the Xe<sup>9+</sup> ions about different *l*, at *n*=5. Figure 2 shows the <Ar<sub>*i',i0*</sub> \* *g<sub>i'</sub>* / *g*<sub>0</sub>> values of the line from radiative decay from Xe<sup>10+</sup> (4d<sup>7</sup> 5p<sup>1</sup>) to Xe<sup>10+</sup> (4d<sup>8</sup>).

Through figures, we know that the dielectronic satellite lines of Xe<sup>9+</sup> ions appear near the radiative decay lines of Xe<sup>10+</sup> ions and the dielectronic satellite lines have values that can not ignore at low temperate. For identification of the spectrum from the highly charged ions, we will study more detailed about the direct excitation of the parent ions.

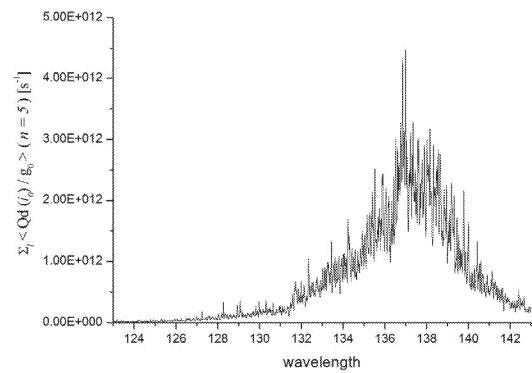


Fig. 1. Satellite line of the transition  
[ Xe<sup>9+</sup> (4d<sup>7</sup> 5p<sup>1</sup> nl) → Xe<sup>9+</sup> (4d<sup>8</sup> nl) + ħω ]

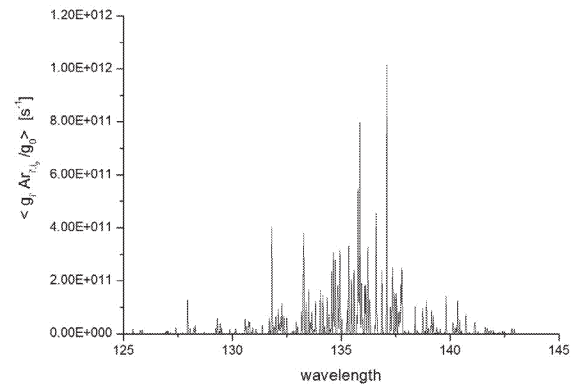


Fig. 2. Radiative line of the transition  
[ Xe<sup>10+</sup> (4d<sup>7</sup> 5p<sup>1</sup>) → Xe<sup>10+</sup> (4d<sup>8</sup>) + ħω ]

#### Reference

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